

## **1** Introduction

Recent developments in computational seismology model parameter. We use the spectral-element allow for highly accurate modelling of wave propa- method for the simulation of 3D wave propagation gation in strongly heterogeneous media. This en- through heterogeneous Earth models. An objectables us to use full seismograms as a data source based parametrisation of the Earth, motivated by for tomographic inversions. The classical approach the geological structure of western and central Austo seismic tomography, however, consists of find- tralia, enables us to limit the dimensionality of the ing one single velocity model of the Earth's interior model space and to test hypothesis efficiently. that minimises the misfit between simulated and ob- We apply our methodology to the Australian contiserved seismograms. It does thus not account for nental lithosphere. This is intended to answer the the possible existence of multiple solutions that ex- following questions: (1) How good is the resolution plain the data equally well, and it does not provide of the tomographic models? (2) Is there reliable ininformation on the reliability of a model. formation on density variations in the upper mantle We circumvent this issue by making use of a Monte contained in the waveforms? (3) How robust is the Carlo optimisation method based on a Bayesian frequently inferred low-velocity layer around 150 km statistical framework. This leads to an ensem- depth beneath Proterozoic Australia? ble of models and to an error estimate for each

### 2 Geological structure

- Identification of two large-scale (≥ 1500 km)
  structural elements in continental Australia: the Archean cratons in the west and the predominantly Proterozoic units in the centre.
- Both Archean and Proterozoic lithosphere have been imaged consistently in several recent studies ([1, 10, 4, 5, 2, 3]). Significant differences between the tomographic images are limited to length scales below about 1500 km.
- Archean lithosphere is marked by anomalously high S wave speeds reaching +8% with respect to the radial average on a 1500 km length scale.



## **3** Object-based model construction

a) Definition of four basis functions  $b_i$  according to the geological structures in the study area in two layers from 0 to 150km and from 150km to 230km, respectively.



**b)** 3D Earth models  $m(\mathbf{x})$  are constructed as filtered superpositions of weighted basis functions  $b_i$ , which take the value one within and zero outside the object. The objects are surrounded by a tomographic Australia model from Fichtner et al. (2009).





 Elastic properties of the Proterozoic lithosphere are strongly depth-dependent with a structural boundary located at around 150 km depth.

**Figure:** Major surface geologic features in the study area. Adapted from Myers et al. (1996).

AB - Amadeus Basin, ARB - Arunta Block, BH - Broken Hill Block, CU - Curnamona Block, GB - Georgina Basin, GC - Gawler Craton, GI - Georgetown Inlier, KB - Kimberley Block, LFB - Lachlan Fold Belt, MAB - McArthur Basin, MB - Musgrave Block, MII - Mount Isa Inlier, NB - Ngalia Basin, NFB - New England Fold Belt, NVP - Newer Volcanic Province, PB - Pilbara Block, PCI - Pine Creek Inlier, YB - Yilgarn Block

# 4 Data & Forward Modelling



#### **Observed data:**

- primarily sensitive to the regions covered by the basis functions
- 338 vertical-component seismograms recorded at 30 different stations

#### Preprocessing:

- two frequency ranges from 60s to 200s and from 130s to 200s, in order to focus the sensitivity on the depth range of interest
- selection of the surface wave part by manually tapering each recording

**Forward problem:** We use a spectral element method as described by Fichtner et al. (2009) to simulate wave propagation in a spherical section ranging from  $7.5^{\circ}$  N to  $50^{\circ}$  S, from  $105^{\circ}$  E to  $160^{\circ}$  E and from the surface to 1461 km in depth.

- 25200 hexahedral elements with a size of approx.  $1.6^{\circ} \times 1.6^{\circ} \times 73$ km
- approximation of the wavefield by 4th order Lagrange polynomials collocated at the Gauss-Lobatto-Legendre points



## 5 Probabilistic Inversion using the Neighbourhood Algorithm

Variation of P velocity-  $\delta v_p$ , S velocity-  $\delta v_s$  and where  $u_0(t)$  and u(t) denote the observed and syndensity-perturbations  $\delta \rho$  with respect to isotropic thetic seismograms. A cumulative misfit  $E(\mathbf{m})$  for PREM in every object  $\rightarrow$  **12-dimensional model** model **m** is given by a weighted sum over the total space.

The posterior probability density (PPD) of a model

space vector  $\mathbf{m}$  given a set of observed data  $\mathbf{d}$  is **c)** A natural choice for the **likelihood function** given by  $L(\boldsymbol{m}|\mathbf{d})$  in terms of the cumulative misfit thus is:

(6)

 $\sigma(\boldsymbol{m}|\mathbf{d}) = \frac{1}{\nu} \rho_M(\boldsymbol{m}) L(\boldsymbol{m}|\mathbf{d}).$ (3)

a) Choice of a uniform prior probability density  $\rho_M(\boldsymbol{m})$  within the intervals:

 $\delta v_s \in [0; 0.6 \text{ km s}^{-1}]$  $\delta v_p \in [0; 0.8 \text{ km s}^{-1}]$  $\delta \rho \in [-0.4 \text{ g cm}^{-3}; 0.4 \text{ g cm}^{-3}]$ 

**b)** Objective functional: A per-seismogram misfit is given by the  $L_1$  distance between the  $L_1$ normalised waveforms:

 $E(u_0, u) := ||u_0/||u_0||_1 - u/||u||_1||_1$ , (7) the joint

(4) where the scale parameter  $\lambda = 1/\sigma$  is related to the variance  $\sigma^2$  of the observed seismograms, reflecting (5) the noisy nature of the measurements.

 $L(\boldsymbol{m}) = \lambda^{N_d} \exp\left[-\lambda E(\mathbf{m})\right],$ 

d) Model space sampling: We use the Neighbourhood Algorithm (Sambridge, 1999) to sample the likelihood function quasi-randomly. The resulting ensemble of models gives an approximation of (7) the joint PPD.

## 6 Results I - Posterior Probability Density

After 18 iterations the Neighbourhood Algorithm has generated 5000 models. From the approximate joint PPD marginal posterior distributions are retrieved by numerical integration. We observe three kinds of distributions:

 single-peaked, nearly Gaussian distributions (Archean and Proterozoic top layer S wave speeds)

 nearly exponential distributions (most densities and S wave speeds in the bottom layer) → might occur due to undersampling or an insufficiently large parameter range, these parameters are thus excluded from interpretation

nearly uniform distributions (most P wave speeds)

#### The results confirm

(8)

- high S wave speed perturbations (around  $0.35 \ km/s$ ) in the uppermost 150 km of Archean lithosphere. With 90% confidence values are higher than  $0.1 \ km/s$ ,
- slightly less pronounced S wave speed perturbations in the Proterozoic part, with values higher than  $0.32 \ km/s$  having a probability of less than 10%.



**Figure:** Marginal posterior distributions for top layer S and P wave speeds.



**Figure:** Horizontal slices through the most likely model in the ensemble.

## 7 Results II - Posterior Model Covariance

## 8 Conclusion & Outlook

### **9** References

- Nearly uniform marginals for the P wave speeds indicate that each single parameter does not have a significant influence.
- The posterior covariance matrix, however, suggests a rather strong (anti-) correlation between the P wave speeds.





**Posterior model covariance matrix**. Row-order: Archeantop, -bottom, Proterozoic-top, -bottom.

• The 2D marginal of Archean top layer vs. Archean bottom layer P wave speed (figure on the left) suggests that either both parameters take very high or very low values (green contour 60%, blue contour 90% and red contour 99% confidence) We showed that the methodology of a probabilistic full waveform inversion based on a regionalized parameterization works by inverting surface wave data to infer information about Australian continental lithosphere. Our results are consistent with recent tomographic studies. A future goal may be to relate the models endowed with error bars to results from mineralogical studies.

However, we note that we defer the interpretation of the density marginals, since we experienced problems possibly due to under-sampling. We note that these might be circumvented by using more informative prior distributions and a different initial tuning of the sampling algorithm.

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